(B)



<u>Context</u>

Jason-1 and TOPEX GDRs have been reprocessed by JPL from Jason-1 cycles 1 to 21 (T/P cycles 344 to 364). During this period, Jason-1 (J1) and TOPEX flew on the same track. The objective of this study is to check the impact of the reprocessing on the sea surface height consistency between J1 and TOPEX.



Data and processing

The main evolutions in the reprocessed data (GDR 'B' for J1 and RGDR for TOPEX) are the implementation of a new retracking algorithm (MLE4 for Jason-1 and MLE5 for TOPEX), a new precise orbit based on a GRACE gravity model and new geophysical corrections (tidal models, MOG2D, Sea State Bias). For comparison between GDR "A" and GDR "B" for J1, please refer to poster (A) Performance assessment J1 Gdr 'A' /Gdr 'B'. In order to get a significant data set, the following statistics are computed over the 21 cycles, excluding cycles TOPEX 361 (J1 18), which was not reprocessed and TOPEX 362, which was almost completely edited by the Cal/Val processing due to spurious altimeter parameters.

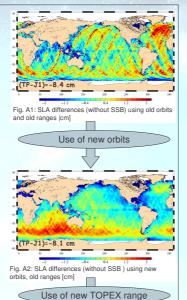
A) SLA consistency : Impact of new orbit

SLA differences between J1 and TOPEX are plotted (figure A1) using former orbits and ranges from GDR 'A' for J1 and MGDR for TOPEX. Neither geophysical corrections nor SSB correction were applied for both satellites. Large structures of negative and positive differences are visible, as well as orbit

Using the new orbits (GRACE family) provided by the GDR 'B' for J1 and RGDR TOPÉX, removes trackiness and decreases the particular pattern in North Atlantic (figure A2).

Thanks to the new orbits, large structures are detected in Indian ocean and close to of the shores. Some part these correspond SSB discrepancies to differences between the two missions.

Besides, a thick equatorial band is evidenced on figure A2 with negative differences. This is due to the ascending and descending SLA differences between J1 and TOPEX showing a large hemispheric signal (see figure B4).

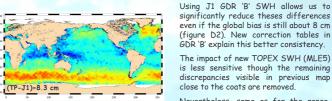


D) SWH consistency

Figure D1: SWH differ TOPEX MGDR [cm]

Figure D2: SWH differences for J1 GDR B and TOPEX MGDR [cm]

Mean differences between TOPEX MGDR and J1 GDR 'A' SWH estimates (figure D1) was about 8 cm, but this bias was higher for strong waves and smaller for small waves



es for J1 GDR A and

significantly reduce theses differences even if the global bias is still about 8 cm (figure D2). New correction tables in GDR 'B' explain this better consistency. The impact of new TOPEX SWH (MLE5)

is less sensitive though the remaining discrepancies visible in previous map close to the coats are removed.

Nevertheless, same as for the range, when separating ascending and when separating ascending and descending passes, large hemispheric biases appear (not shown here) due to the TOPEX SWH.

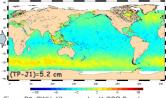


Figure D3: SWH differences for J1 GDR B and TOPEX RGDR [cm]

B) SLA consistency: Impact of new range

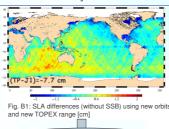
When using the new range (from MLE5 algorithm) for TOPEX, the patch in Indian $\frac{1}{2}$ ocean is strongly reduced (figure B1). Jason-1 and TOPEX SSB are probably more homogeneous from now on (see dedicated S. Labroue's talk).

Using the new MLE4 range for Jason (Figure B2) has weak impact on the mean differences, even if the consistency is slightly better in the Indian Ocean

Nevertheless a great hemispheric bias (between -2 cm and +2 cm) is highlighted when separating the ascending and descending passes (figure B4):

- This bias is mainly due to TOPEX data. It was present on TOPEX M-GDR data alone (figure B3, left), but it is greater using new range (MLE5 from RGDR) as shown at the TOPEX crossovers (figure B3, right). This needs more investigation

- To a lower extent, such a signal is also visible at Jason-1 crossovers in the GDR 'B' (see CalVal poster 'A') probably due to time tag bias. But it is much weaker than for



Use of new J1 range

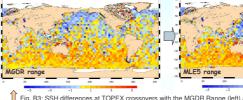
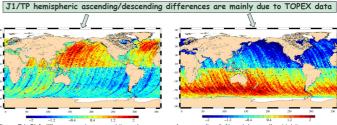


Fig. B3: SSH differences at TOPEX crossovers with the MGDR Range (left) and MLE5 range (right). ${\small \textbf{TOPEX MLE5 range increases}} \ \ {\small \textbf{the hemispheric ascending/descending differences}}$

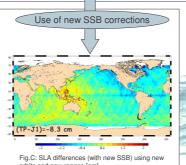
Fig. B2: SLA differences (without SSB) using new orbits new ranges [cm]



ferences using new orbits and new ranges for ascending (left) and descending (right) passes

C) SLA consistency: Impact of new SSB

New SSB corrections have been computed on cycles 1-21 for J1 using GDR 'B' and for TOPEX using RGDR, with the collinear method. For more details see dedicated talk by S. Labroue. These new TOPEX and J1 SSB models are now much closer than before. When applying them in the SLA calculation in addition to the new orbits and the new ranges (figure C), the discrepancies between J1 and T/P are reduced. However, an East/West patch (< 1cm) remains, but it is not correlated with SWH.



. B5: STD of SLA

STD(TP-J1)= 4.11 cm

Impact of new range in standard <u>deviation differences</u>

Thanks to the MLE4 retracking, the STD differences are dropped from 4.42 cm (figure B5) to 4.11 cm (figure B6) which is consistent with results shown in CaVal poster 'A'. Use of the T/P MLE5 range (figure B7) slightly increases the STD differences.

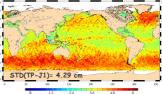


Fig. B6: STD of SLA differences using new range [cm] Fig. B7: STD of SLA differences using new range [cm]