SSH and velocity frequency and wavenumber spectra in the Atlantic Ocean, estimated from altimetry and a hierarchy of numerical simulations

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Model velocity spectra as for SSH, improve with increasing resolution. Most differences either in shape and/or magnitude are located in the tropics.



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With increasing resolution of the model, model spectra closer resemble altimetric results. However, even the highest resolution available spectra fails to describe the same spectral density contained in TJJ at high frequencies, though the spectral slopes tend to be qualitatively close.



There is a clear shift towards smaller scales and higher amplitudes as the model resolution increases rendering the spectra much closer to TJJ. However, an obvious unexplained wedge remains between altimetry and model results on high wavenumbers where the model spectra have steeper slopes than TJJ.

Frequency Spectra – Velocity

UH

Wavenumber Spectra – Velocity



There is a good agreement both in terms of amplitude and shape between Aviso and 8 km velocities, especially in mid and high latitudes. The 16 and 32 km have steeper slopes than Aviso in high frequencies. As seen for SSH, also for velocities with the increase of model resolution, the spectra shape draws closer to the one derived from Aviso.

Methodology The altimeter data were interpolated in time and space. A windowed FFT method was used to compute the frequency and wavenumber spectra. Before the computation of the frequency spectra any remaining trend was removed. Frequency spectra were computed for every along-track position and then averaged globally or regionally. For the computation of wavenumber spectra of SSH the methodology described by Scharffenberg and Stammer [2011] was followed. The MITgcm and AVISO velocity spectra were computed on the grid.

Reference: Scharffenberg, M. and Stammer, D. (2011). Statistical parameters of the geostrophic ocean flow field, estimated by jason-1-topex/poseidon tandem mission. J. Geophys. Res., 116(C12011):1--14, doi:10.1029/2011JC007376.